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Abstract

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Williamson–Hadamard spreading sequences for DS-CDMA applications

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Summary

Orthogonal bipolar spreading sequences are used in direct sequence code division multiple access (DS-CDMA) systems for both spectrum spreading and channel separation. The most commonly used sequences are Walsh–Hadamard sequences of lengths being an integer power of 2. A construction based on Williamson’s arrays leading to sequences of lengths $N \equiv 4 \pmod{8}$ is presented in the paper. Aperiodic correlation characteristics, for example sequence sets of lengths 12–252 are presented. The correlation properties of the sequence sets are later improved using a diagonal modification technique. Copyright © 2003 John Wiley & Sons, Ltd.

KEY WORDS: Williamson’s arrays; spreading sequences; code division multiple access; correlation functions

1. Introduction

Orthogonal bipolar sequences are of a great practical interest for the current and future direct sequence (DS) code division multiple access (CDMA) systems where the orthogonality principle can be used for channels separation, e.g. Reference [1]. The most commonly used sets of bipolar sequences are Walsh–Hadamard sequences [2], as they are easy to generate and simple to implement. However, they exist only for sequence lengths being an integer power of 2, which can be a limiting factor in some applications. In the paper, we describe a technique to generate sets of bipolar sequences of order $N \equiv 4 \pmod{8}$ based on a Williamson’s construction [3]. The resultant Williamson–Hadamard sequences possess very good autocorrelation properties that make them amenable to synchronization requirements.

It is well known, e.g. Reference [4–6], that if the sequences have good aperiodic cross-correlation properties, the transmission performance can be im-

proved for those CDMA systems where different propagation delays exist. Wysocki and Wysocki in Reference [7] proposed a technique to modify bipolar Walsh–Hadamard sequences to achieve changes in their correlation characteristics without compromising orthogonality. In this paper, we apply the same technique to improve cross-correlation properties of Williamson–Hadamard sequences. As it is always the case, the improvement is achieved at the expense of slightly worsening the autocorrelation properties. However, the overall autocorrelation properties of the modified sequence sets are still significantly better than those of Walsh–Hadamard sequences of comparable lengths.

The paper is organized as follows. In Section 2, we introduce principles of constructing Hadamard matrices using Williamson’s arrays and provide a list of some possible seed sequences of lengths 3–63 to construct Williamson–Hadamard matrices of orders 12–252. Section 3 introduces some correlation measures that can be used to compare different sets of

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spreading and show the values of those parameters for the sequence sets derived utilizing the seed sequences listed in Section 2. In Section 4, we briefly describe the method used to modify correlation characteristics of sequence sets and show the results when applied to Williamson–Hadamard sequences. Section 5 concludes the paper.

2. Williamson–Hadamard Construction

A Hadamard matrix \mathbf{H} of order n has elements ± 1 and satisfies $\mathbf{H}\mathbf{H}^T = n\mathbf{I}_n$. The order of a Hadamard matrix is 1, 2 or $n \equiv (0 \pmod 4)$ and the first unsolved case is order 428. We briefly describe the theory of Williamson’s construction below. Previous computer searches for Hadamard matrices using Williamson’s condition are described in Section 2.1.

Theorem 1 (Williamson [8]): *Suppose there exist four symmetric $(1, -1)$ matrices $\mathbf{A}, \mathbf{B}, \mathbf{C}, \mathbf{D}$ of order n which satisfy*

$$\mathbf{X}\mathbf{Y}^T = \mathbf{Y}\mathbf{X}^T, \quad \mathbf{X}, \mathbf{Y} \in \{\mathbf{A}, \mathbf{B}, \mathbf{C}, \mathbf{D}\}.$$

Further, suppose

$$\mathbf{A}\mathbf{A}^T + \mathbf{B}\mathbf{B}^T + \mathbf{C}\mathbf{C}^T + \mathbf{D}\mathbf{D}^T = 4n\mathbf{I}_n. \quad (1)$$

Then

$$\mathbf{H} = \begin{bmatrix} \mathbf{A} & \mathbf{B} & \mathbf{C} & \mathbf{D} \\ -\mathbf{B} & \mathbf{A} & -\mathbf{D} & \mathbf{C} \\ -\mathbf{C} & \mathbf{D} & \mathbf{A} & -\mathbf{B} \\ -\mathbf{D} & -\mathbf{C} & \mathbf{B} & \mathbf{A} \end{bmatrix} \quad (2)$$

is a Hadamard matrix of order $4n$ constructed from a Williamson array.

Let the matrix \mathbf{T} given below be called the shift matrix:

$$\mathbf{T} = \begin{bmatrix} 0 & 1 & 0 & \dots & 0 \\ 0 & 0 & 1 & \dots & 0 \\ \vdots & \vdots & \vdots & \dots & \vdots \\ 0 & 0 & 0 & \dots & 1 \\ 1 & 0 & 0 & \dots & 0 \end{bmatrix} \quad (3)$$

and note that

$$\mathbf{T}^n = \mathbf{I}, \quad (\mathbf{T}^i)^T = \mathbf{T}^{n-i}. \quad (4)$$

If n is odd, \mathbf{T} is the matrix representation of the n th root of unity ω , $\omega^n = 1$.

Let

$$\begin{cases} \mathbf{A} = \sum_{i=0}^{n-1} a_i \mathbf{T}^i, & a_i = \pm 1, & a_{n-i} = a_i \\ \mathbf{B} = \sum_{i=0}^{n-1} b_i \mathbf{T}^i, & b_i = \pm 1, & b_{n-i} = b_i \\ \mathbf{C} = \sum_{i=0}^{n-1} c_i \mathbf{T}^i, & c_i = \pm 1, & c_{n-i} = c_i \\ \mathbf{D} = \sum_{i=0}^{n-1} d_i \mathbf{T}^i, & d_i = \pm 1, & d_{n-i} = d_i. \end{cases} \quad (5)$$

Then matrices $\mathbf{A}, \mathbf{B}, \mathbf{C}, \mathbf{D}$ may be represented as polynomials. The requirement that $x_{n-i} = x_i, x \in \{a, b, c, d\}$, forces the matrices $\mathbf{A}, \mathbf{B}, \mathbf{C}, \mathbf{D}$ to be symmetric. Hereafter, we will refer to the sequences a, b, c, d , as the seed sequences.

Since $\mathbf{A}, \mathbf{B}, \mathbf{C}, \mathbf{D}$ are symmetric, Equation (1) becomes

$$\mathbf{A}^2 + \mathbf{B}^2 + \mathbf{C}^2 + \mathbf{D}^2 = 4n\mathbf{I}_n,$$

and the relation $\mathbf{X}\mathbf{Y}^T = \mathbf{Y}\mathbf{X}^T$ becomes $\mathbf{X}\mathbf{Y} = \mathbf{Y}\mathbf{X}$ which is true for polynomials.

Definition 1: *Williamson matrices are $(1, -1)$ symmetric circulant matrices.*

As a consequence of being symmetric and circulant, they commute in pairs.

The following theorem of Williamson has been used as the motivator for search algorithm:

Theorem 2 (Williamson [8]): *If there exist solutions to the equations*

$$\mu_i = 1 + 2 \sum_{j=1}^s t_{ij} (\omega^j + \omega^{n-j}), \quad i = 1, 2, 3, 4, \quad (6)$$

where $s = \frac{1}{2}(n - 1)$, ω is a n th root of unity, exactly one of $t_{1j}, t_{2j}, t_{3j}, t_{4j}$, is nonzero and equals ± 1 for each $1 \leq j \leq s$, and

$$\mu_1^2 + \mu_2^2 + \mu_3^2 + \mu_4^2 = 4n,$$

then there exist solutions to the equations:

$$\begin{cases} \mathbf{A} = \sum_{i=0}^{n-1} a_i \mathbf{T}^i, & a_0 = 1, & a_i = a_{n-i} = \pm 1 \\ \mathbf{B} = \sum_{i=0}^{n-1} b_i \mathbf{T}^i, & b_0 = 1, & b_i = b_{n-i} = \pm 1 \\ \mathbf{C} = \sum_{i=0}^{n-1} c_i \mathbf{T}^i, & c_0 = 1, & c_i = c_{n-i} = \pm 1 \\ \mathbf{D} = \sum_{i=0}^{n-1} d_i \mathbf{T}^i, & d_0 = 1, & d_i = d_{n-i} = \pm 1. \end{cases} \quad (7)$$

That is, there exists a Hadamard matrix of order $4n$.

In matrix form, $\omega^j + \omega^{n-j}$ is represented as $\mathbf{T}^j + \mathbf{T}^{n-j}$. Since these are symmetric, we write

$$\omega = \omega^j + \omega^{n-j}.$$

Remark 1: The solutions for Equation (6) are independent of the particular root ω , so if n as defined by Equation (1) is prime, ω can be chosen so that the first μ having any ω_j assigned has ω_1 . Since the equations are true for all roots of unity ω , they are also true for $\omega = 1$.

2.1. Results From Previous Searches

In many cases, complete searches have been conducted for Hadamard matrices of Williamson type. Searches have also been conducted for special classes of Williamson type Hadamard matrices. Furthermore, an infinite class of such matrices is known and will also be discussed briefly.

- Williamson [8] used algebraic and number theoretic results to simplify his first searches by hand in 1944.
- Baumert and Hall [9] report results of a complete search for orders $4t$, t odd and $3 \leq t \leq 23$. Some incomplete results for higher orders are also given.
- Sawade [10] reports results of a complete search for orders $4t$, $t = 25, 27$. The results for $t = 25$ were later demonstrated to be incomplete by Djokovic [11].
- Djokovic [12] reports results of a complete search for orders $4t$, $t = 29, 31$. Only a single non-equivalent solution was found for $t = 29$ and is equivalent to an earlier result of Baumert [13].
- Koukouvinos and Kounias [14, 15] report results of a complete search for order $4t$, $t = 33$ and 39. These results were complete search for order $4t$, $t = 33$ and 39. These results were later demonstrated to be incomplete by Djokovic [16].
- Djokovic [16] reports results of a complete search for orders $4t$, $t = 33, 35, 39$.
- Djokovic [11] reports results of a complete search for orders $4t$, $t = 25, 37$. This extends results obtained by Sawade [10] for $t = 25$ and, for $t = 37$, by Williamson [9] and later Yamada [17] for a special class of matrices.
- Horton *et al.* [18] report results of a complete search for orders $4t$, t odd and $25 \leq t \leq 37$. No new results were found, confirming existence results.

An infinite family of Hadamard matrices of Williamson type has been proved to exist under certain conditions [19, 20]:

Theorem 3 (Williamson [9]): *If q is a prime power, $q \equiv 1 \pmod{4}$, $q + 1 = 2t$, then there exists a Williamson matrix of order $4t$; we have $\mathbf{C} = \mathbf{D}$, and \mathbf{A} and \mathbf{B} differ only on the main diagonal.*

This theorem gives examples of Hadamard matrices of Williamson type for orders $4t$, $t = 31, 37, 41, 45, 49, 51, 55, \dots$, for example.

Yamada [17] has searched for Hadamard matrices of Williamson type, with certain restrictions. These matrices are referred to as Williamson type j matrices. The Williamson equation for such matrices, of order $4n$ is:

$$4n = \left(1 - 2 \sum_{s \in A} c_s \omega_s\right)^2 + \left(1 - 2 \sum_{s \in A} c_s \omega_{sj}\right)^2 + \left(1 - 2 \sum_{s \in B} c_s \omega_s\right)^2 + \left(1 - 2 \sum_{s \in B} c_s \omega_{sj}\right)^2, \tag{8}$$

where $c_s, d_s = \pm 1$, $\omega_s = \omega^s + \omega^{-s}$, $\omega^n = 1$, $j^2 \equiv -1 \pmod{n}$, $\mathbf{A}, \mathbf{B}, j\mathbf{A}, j\mathbf{B}$ is a partition of $\{1, 2, \dots, (n-1)/2\}$. Such a j exists, if and only if, all prime divisors of n are $\equiv 1 \pmod{4}$. This led to some new results for $n = 29, 37, 41$. A summary of the presently known results can be found in Horton *et al.* [18].

In Table 1, we list some of the seed sequences that can be used to construct Walsh–Hadamard sequence sets for DS-CDMA applications. This list includes just a single quadruple of sequences for a given length.

3. Correlation Measures

It is well known (e.g. References [4, 5]) that the level of multi-access interference and synchronization amenability depend on the cross-correlations between the sequences and the autocorrelation functions of the sequences respectively. In this section, we introduce some of the quantitative measures based on the aperiodic correlation functions that can be used to compare the sequence sets from the viewpoint of their usefulness in DS-CDMA systems. Then, we present the computed values of these measures for the Williamson–Hadamard sequences created using formula (2) from the seed sequences listed in Table I.

For general polyphase sequences $\{s_n^{(i)}\}$ and $\{s_n^{(l)}\}$ of length N , the discrete aperiodic correlation function is defined as [5,22]:

$$c_{i,l}(\tau) = \begin{cases} \frac{1}{N} \sum_{n=0}^{N-1-\tau} s_n^{(i)} [s_{n+\tau}^{(l)}]^*, & 0 \leq \tau \leq N-1 \\ \frac{1}{N} \sum_{n=0}^{N-1+\tau} s_{n-\tau}^{(i)} [s_n^{(l)}]^*, & 1-N \leq \tau < 0 \\ 0, & |\tau| \geq N, \end{cases} \tag{9}$$

Table I. List of the seed sequences for Williamson–Hadamard sequences of length 12–252.

<i>N</i>	Seed sequences
12	A +++
	B +--
	C +- -
	D +--
20	A +-----
	B +-----
	C ++--++
	D +-+++
28	A +-----
	B ++-----+
	C +-+-----
	D +-----
36	A ++-----+
	B ++-----+
	C +-----+
	D +-----
44	A +++-----+
	B +++-----+
	C +++-----+
	D +++-----+
52	A +++-----+
	B +++-----+
	C +++-----+
	D +++-----+
60	A +++-----+
	B +++-----+
	C +++-----+
	D +++-----+
68	A +++-----+
	B +++-----+
	C +++-----+
	D +++-----+
76	A +++-----+
	B +++-----+
	C +++-----+
	D +++-----+
84	A +++-----+
	B +++-----+
	C +++-----+
	D +++-----+
92	A +++-----+
	B +++-----+
	C +++-----+
	D +++-----+
100	A +++-----+
	B +++-----+
	C +++-----+
	D +++-----+
108	A +++-----+
	B +++-----+
	C +++-----+
	D +++-----+
116	A +++-----+
	B +++-----+
	C +++-----+
	D +++-----+
124	A +++-----+
	B +++-----+
	C +++-----+
	D +++-----+
132	A +++-----+
	B +++-----+
	C +++-----+
	D +++-----+

Continues

Table I. Continued

148	A	+--++++++-----+--+-----+--+-----+++++---
	B	++--++--+-----+--+--+-----+--+--+-----+
	C	+++++-----+--+--+-----+--+--+-----+
	D	+--++++++-----+--+--+-----+--+--+-----+
156	A	+++++-----+--+--+-----+--+--+-----+
	B	+++++-----+--+--+-----+--+--+-----+
	C	+++++-----+--+--+-----+--+--+-----+
	D	+--++++++-----+--+--+-----+--+--+-----+
164	A	+--++++++-----+--+--+-----+--+--+-----+
	B	+--++++++-----+--+--+-----+--+--+-----+
	C	+--++++++-----+--+--+-----+--+--+-----+
	D	+--++++++-----+--+--+-----+--+--+-----+
172	A	+--++++++-----+--+--+-----+--+--+-----+
	B	+++++-----+--+--+-----+--+--+-----+
	C	+++++-----+--+--+-----+--+--+-----+
	D	+++++-----+--+--+-----+--+--+-----+
180	A	+++++-----+--+--+-----+--+--+-----+
	B	+++++-----+--+--+-----+--+--+-----+
	C	+++++-----+--+--+-----+--+--+-----+
	D	+++++-----+--+--+-----+--+--+-----+
196	A	+++++-----+--+--+-----+--+--+-----+
	B	+++++-----+--+--+-----+--+--+-----+
	C	+++++-----+--+--+-----+--+--+-----+
	D	+++++-----+--+--+-----+--+--+-----+
220	A	+--++++++-----+--+--+-----+--+--+-----+
	B	+++++-----+--+--+-----+--+--+-----+
	C	+++++-----+--+--+-----+--+--+-----+
	D	+++++-----+--+--+-----+--+--+-----+
228	A	+--++++++-----+--+--+-----+--+--+-----+
	B	+++++-----+--+--+-----+--+--+-----+
	C	+++++-----+--+--+-----+--+--+-----+
	D	+++++-----+--+--+-----+--+--+-----+
244	A	+++++-----+--+--+-----+--+--+-----+
	B	+++++-----+--+--+-----+--+--+-----+
	C	+++++-----+--+--+-----+--+--+-----+
	D	+++++-----+--+--+-----+--+--+-----+
252	A	+++++-----+--+--+-----+--+--+-----+
	B	+++++-----+--+--+-----+--+--+-----+
	C	+++++-----+--+--+-----+--+--+-----+
	D	+++++-----+--+--+-----+--+--+-----+

where $[\bullet]^*$ denotes a complex conjugate operation. When $\{s_n^{(i)}\} = \{s_n^{(l)}\}$, Equation (9) defines the discrete aperiodic autocorrelation function.

In order to evaluate the performance of a whole set of M spreading sequences, the average mean-square value of cross-correlation for all sequences in the set, denoted by R_{CC} , was introduced by Oppermann and Vucetic [5] as a measure of the set cross-correlation performance:

$$R_{CC} = \frac{1}{M(M-1)} \sum_{i=1}^M \sum_{\substack{k=1 \\ k \neq i}}^M \sum_{\tau=1-N}^{N-1} |c_{i,k}(\tau)|^2. \quad (10)$$

A similar measure, denoted by R_{AC} , was introduced in Reference [5] for comparing the autocorrelation performance:

$$R_{AC} = \frac{1}{M} \sum_{i=1}^M \sum_{\substack{\tau=1-N \\ \tau \neq 0}}^{N-1} |c_{i,i}(\tau)|^2. \quad (11)$$

The measure defined by Equation (11) allows for comparison of the autocorrelation properties of the set of spreading sequences on the same basis as the cross-correlation properties.

The measures defined by Equations (10) and (11) are very useful for large sets of sequences and large number of active users, when the constellation of interferers (i.e. relative delays among the active users and the spreading sequences used) changes randomly for every transmitted information symbol. However, for a more static situation, when the constellation of interferers stays constant for the duration of many information symbols, it is also important to consider the worst-case scenarios. This can be accounted for by analyzing the maximum value of peaks in the aperiodic cross-correlation functions over the whole set of sequences and in the aperiodic autocorrelation function for $\tau \neq 0$. Hence, we introduce here two

Table II. Correlation parameters of the Williamson–Hadamard sequences of length $N=12$ –252 obtained from the seed sequences listed in Table I.

N	R_{CC}	R_{AC}	C_{\max}	A_{\max}
12	0.9082	1.00930	0.9167	0.6667
20	0.9675	0.61800	0.9500	0.4000
28	0.9813	0.50510	0.9643	0.2857
36	0.9818	0.63820	0.9722	0.3333
44	0.9874	0.54040	0.9773	0.2727
52	0.9854	0.74700	0.9808	0.3077
60	0.9894	0.62730	0.9833	0.4000
68	0.9898	0.68330	0.9853	0.2941
76	0.9916	0.62770	0.9868	0.3158
84	0.99355	0.53507	0.9881	0.26190
92	0.99246	0.68655	0.98913	0.27174
100	0.99300	0.69333	0.9900	0.26000
108	0.99379	0.66439	0.99074	0.22222
116	0.99461	0.62025	0.99138	0.23276
124	0.99323	0.83231	0.99194	0.23387
132	0.99530	0.61556	0.99242	0.21970
148	1.01010	0.64471	0.99324	0.22297
156	0.99516	0.74966	0.99359	0.19231
164	0.99656	0.56130	0.99390	0.23780
172	0.99641	0.61407	0.99419	0.19186
180	0.99668	0.59480	0.99444	0.21667
196	0.99714	0.55864	0.99490	0.24490
220	0.99736	0.57722	0.99545	0.20909
228	0.99760	0.54495	0.99561	0.27632
244	0.99769	0.56104	0.9959	0.19262
252	0.99762	0.5984	0.99603	0.21825

additional measures to compare the spreading sequence sets:

- Maximum value of the aperiodic cross-correlation functions C_{\max}

$$c_{\max}(\tau) = \max_{\substack{i=1,\dots,M \\ k=1,\dots,M \\ i \neq k}} |c_{i,k}(\tau)|;$$

$$\tau = (-N + 1), \dots, (N - 1)$$

$$C_{\max} = \max_{\tau} \{c_{\max}(\tau)\}. \quad (12)$$

- Maximum value of the off-peak aperiodic autocorrelation functions

$$a_{\max}(\tau) = \max_{k=1,\dots,M} |c_{k,k}(\tau)|;$$

$$A_{\max} = \max_{\tau \neq 0} \{a_{\max}(\tau)\}. \quad (13)$$

In Table II, we list the correlation parameters R_{CC} , R_{AC} , C_{\max} and A_{\max} , computed for the Williamson–Hadamard sequences of lengths 12–252 created using the seed sequences listed in Table I. In Figures 1 and 2, we show the typical behavior of the peaks in the

aperiodic cross-correlation functions $c_{\max}(\tau)$ and peaks in the aperiodic autocorrelation functions $a_{\max}(\tau)$ respectively.

4. Modification Method

Further improvement to the values of correlation parameters of the sequence sets based on Williamson–Hadamard matrices can be obtained using the method introduced in Reference [7] for Walsh–Hadamard sequences. That method is based on the fact that for a matrix \mathbf{H} to be orthogonal, it must fulfill the condition $\mathbf{H}\mathbf{H}^T = N\mathbf{I}$, where \mathbf{H}^T is the transposed Hadamard matrix of order N and \mathbf{I} is the $N \times N$ unity matrix. In the case of Williamson–Hadamard matrices, we have $N=4n$. The modification is achieved by taking another orthogonal $N \times N$ matrix \mathbf{D}_N , and the new set of sequences is based on a matrix \mathbf{W}_N , given by:

$$\mathbf{W}_N = \mathbf{H}\mathbf{D}_N. \quad (14)$$

Of course, the matrix \mathbf{W}_N is also orthogonal [7].

In Reference [7], it has been shown that the correlation properties of the sequences defined by \mathbf{W}_N can be

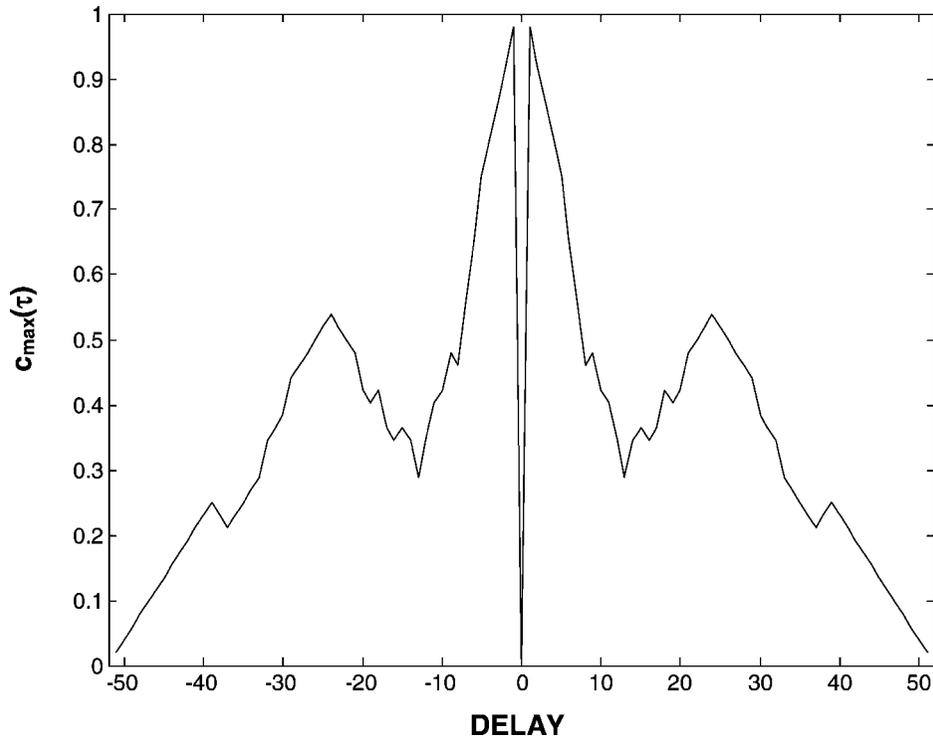


Fig. 1. Plot of the peaks in cross-correlation functions for the Williamson-Hadamard sequence set of order $N=52$.

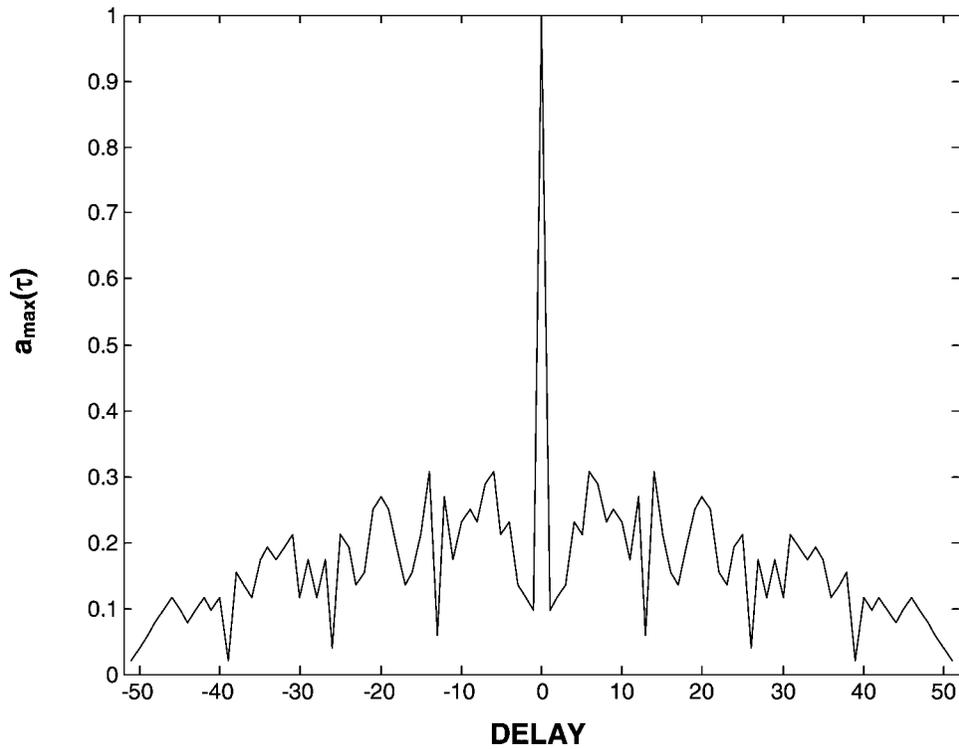


Fig. 2. Plot of the peaks in autocorrelation functions for the Williamson-Hadamard sequence set of order $N=52$.

significantly different than those of the original sequences.

A simple class of orthogonal matrices of any order are diagonal matrices with their elements $d_{i,j}$ fulfilling the condition:

$$|d_{l,m}| = \begin{cases} 0 & \text{for } l \neq m; \\ k & \text{for } l = m; \end{cases} \quad l, m = 1, \dots, N. \quad (15)$$

To preserve the normalization of the sequences, the elements of \mathbf{D}_N , being in general complex numbers, must be of the form:

$$d_{l,m} = \begin{cases} 0 & \text{for } l \neq m; \\ \exp(j\phi_l) & \text{for } l = m; \end{cases} \quad (16)$$

$l, m = 1, \dots, N.$

From the implementation point of view, the best class of sequences are binary sequences.

To find the best possible modifying diagonal matrix \mathbf{D}_N , we can do an exhaustive search of all possible bipolar sequences of length N and choose the one which leads to the best performance of the modified set of sequences. However, this approach is very computationally intensive, and even for a modest value of N , e.g. $N=28$, it is rather impractical. Hence, other search methods, like a random search, must be considered.

By applying a Monte Carlo algorithm [23,24] to $N \geq 20$ and looking for a minimum value of the peaks in the aperiodic cross-correlation functions C_{\max} in 5000 random draws, we have found the sequences listed in Table III for lengths 20–100, and in Table IV, we present the corresponding correlation parameters of the modified sequence sets. In Figures 3 and 4, we present examples of the plots of the peaks in the aperiodic cross-correlation functions $c_{\max}(\tau)$ and peaks in the aperiodic autocorrelation functions $a_{\max}(\tau)$ respectively for the modified sequence sets. There, it is clearly visible that the peaks in the cross-correlation functions are significantly reduced compared to the original sequence sets, shown in the figures by the dotted line. However, this is done on the expense of lifting the peaks in the off-peak autocorrelation functions.

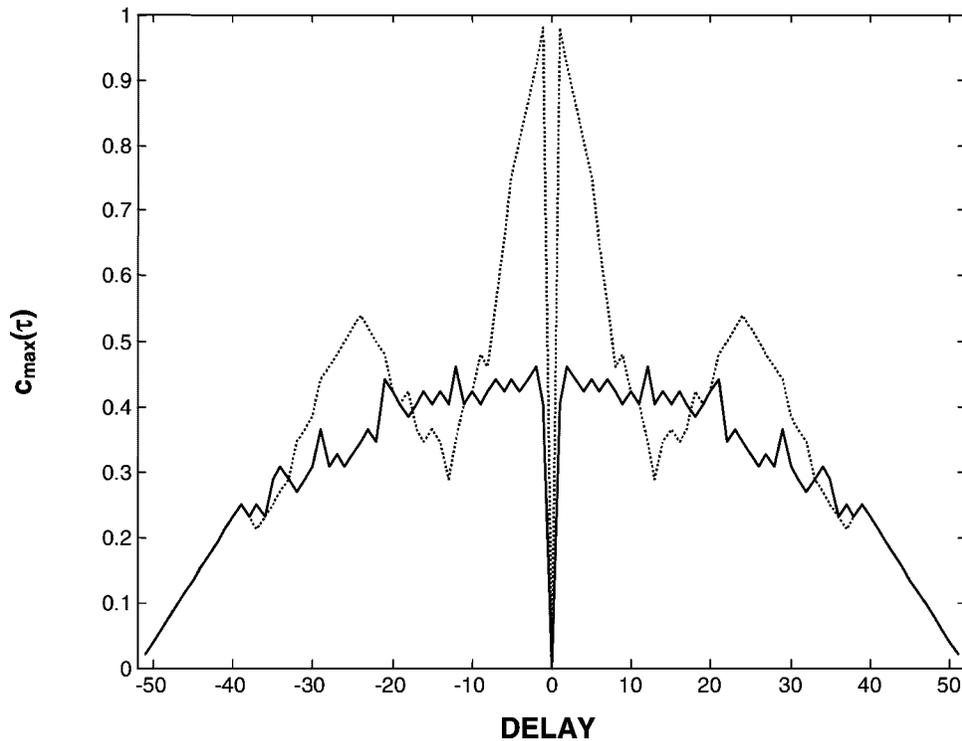
Because of the nonlinear character of the cost function, it is difficult to assess how far the obtained result is from the global minimum without performing the exhaustive search. Calculating the theoretical

Table III. Modifying diagonal sequences for the sequence lengths $N = 20-100$ for the Williamson–Hadamard sequences obtained from the seed sequences listed in Table I.

N	D_N
20	-----++
28	-----++
36	-----++
44	-----++
52b	-----++
60	-----++
68	-----++
76	-----++
84	-----++
92	-----++
100	-----++

Table IV. Correlation parameters of the Williamson–Hadamard sequences of length N modified using a diagonal method to find minimum of AC_{\max} .

N	R_{CC}	R_{AC}	C_{\max}	A_{\max}	B_W	B_L
20	0.9321	1.2900	0.5500	0.6500	0.1562	0.1759
28	0.9664	0.9074	0.5357	0.4643	0.1325	0.1503
36	0.9764	0.8248	0.5000	0.4444	0.1170	0.1333
44	0.9777	0.9604	0.4773	0.5227	0.1060	0.1210
52	0.9804	0.9987	0.4615	0.3654	0.0976	0.1116
60	0.9838	0.9539	0.4500	0.3833	0.0909	0.1041
68	0.9854	0.9795	0.4265	0.4412	0.0854	0.0979
76	0.9877	0.9210	0.4079	0.3026	0.0808	0.0927
84	0.9879	1.0008	0.3929	0.4048	0.0769	0.0883
92	0.9885	1.0451	0.3913	0.3261	0.0735	0.0844
100	0.9902	0.9660	0.3800	0.3400	0.0705	0.0810

Fig. 3. Plot of the peaks in cross-correlation functions for the modified Williamson–Hadamard sequence set of order $N=52$; the dotted line represents the values for the original sequence set.

lower bound for the aperiodic cross-correlation and aperiodic out-of-phase autocorrelation magnitudes can give some insight into this. The best-known bound is given by Welch [25] and states that for any set of M bipolar sequences of length N

$$\max\{C_{\max}, A_{\max}\} \geq \sqrt{\frac{M-1}{2NM-M-1}} = B_W. \quad (17)$$

A more tighter bound was given by Levenshtein [26] and is expressed by:

$$\max\{C_{\max}, A_{\max}\} \geq \sqrt{\frac{(2N^2+1)M-3N^2}{3N^2(MN-1)}} = B_L. \quad (18)$$

The values of both B_W and B_L calculated for the considered values of N are also listed in Table IV.

It must be noted here that both Welch and Levenshtein bounds are derived for sets of bipolar sequences where the condition of orthogonality for perfect

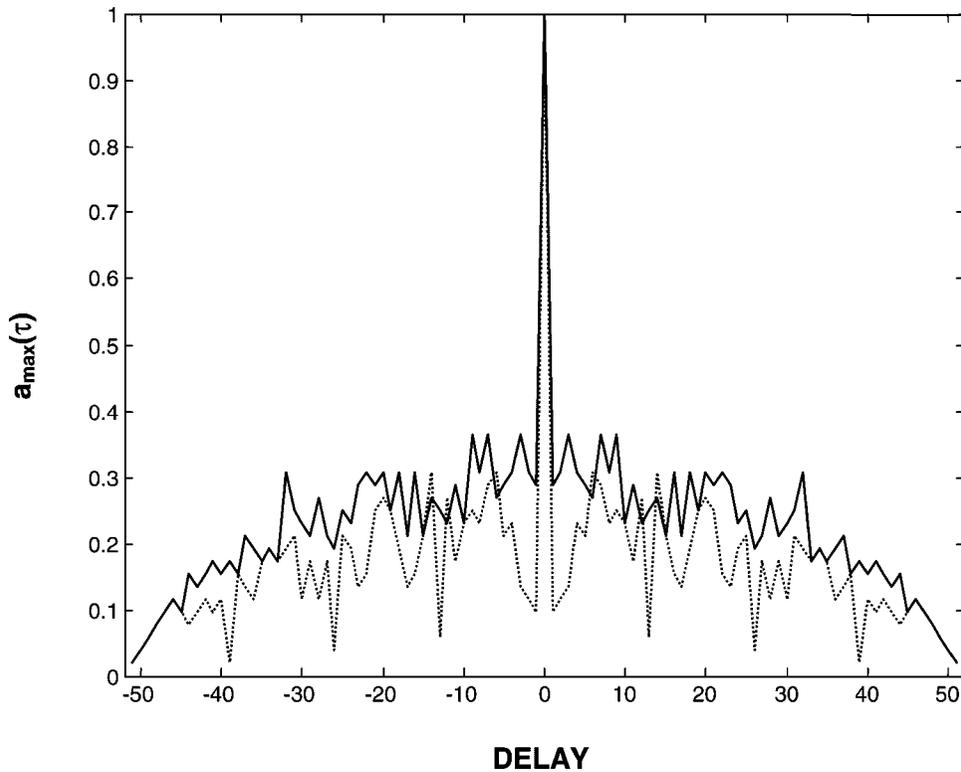


Fig. 4. Plot of the peaks in autocorrelation functions for the modified Williamson–Hadamard sequence set of order $N = 52$; the dotted line represents the values for the original sequence set.

synchronization is not imposed. Hence, one can expect that by introducing the orthogonality condition, the lower bound for the aperiodic cross-correlation and aperiodic out-of-phase autocorrelation magnitudes must be significantly lifted.

5. Conclusions

In this paper, we presented a family of bipolar orthogonal spreading sequences of order $N \equiv 4 \pmod{8}$ based on a Williamson's construction. These sequences possess very good autocorrelation properties that make them amenable to synchronization requirements. Later, we applied a modification technique to improve cross-correlation properties of Williamson–Hadamard sequences without compromising their orthogonality. The improvement was achieved at the expense of slightly worsening the autocorrelation properties. However, the overall autocorrelation properties of the modified sequence sets are still significantly better than those of Walsh–Hadamard sequences of comparable lengths. The proposed family of bipolar spreading sequences can be very useful in those DS-CDMA systems that require spreading different than by a factor being an integer power of 2.

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Authors' Biographies



Dr Seberry was awarded Ph.D. in Computation Mathematics from La Trobe University in 1971. She has subsequently held positions at the Australian National University, The University of Sydney and ADFA, The University of New South Wales. She has published extensively in Discrete Mathematics and is world renowned for her new discoveries on Hadamard Matrices and Statistical Designs. She started

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